



# Spacetime Quantum Geometry

Peter Schupp

Jacobs University Bremen

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## Spacetime quantum geometry

Applying the principles of quantum mechanics to space & time

### Outline

- ▶ spacetime non-commutativity
- ▶ particle physics on non-commutative spaces
- ▶ non-commutative gravity and fuzzy black holes
- ▶ large scale application of small scale math (CMB analysis)
- ▶ higher geometric structures

# Spacetime non-commutativity

## Planck scale quantum geometry



Heuristic argument: quantum + gravity

*“The gravitational field generated by the concentration of energy required to localize an event in spacetime should not be so strong as to hide the event itself to a distant observer.”*

→ fundamental length scale, spacetime uncertainty

$$\Delta x \geq \sqrt{\frac{\hbar G}{c^3}} \approx 1.6 \times 10^{-33} \text{cm}$$

uncertainty principle  $\leftrightarrow$  noncommutative spacetime structure

$$[\hat{x}^i, \hat{x}^j] = i\theta^{ij}(x)$$

## Macroscopic non-commutativity

- ▶ Landau levels

charged particle in a constant magnetic field  $\vec{B} = \nabla \times \vec{A}$ :  
quantize  $\vec{p} = m\dot{\vec{x}} - e\vec{A}$  conjugate to  $\vec{x}$

for  $m \rightarrow 0$  (projection onto 1st Landau level):

$$[\hat{x}^i, \hat{x}^j] = \frac{2i}{eB} \epsilon^{ij} =: \theta^{ij}$$

- ▶ Fractional Quantum Hall Effect

Non-commutative (NC) fluids, NC Chern Simons theory

## Loop quantum gravity, spacetime foam

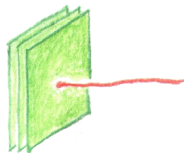
- ▶ non-perturbative framework
- ▶ defines quantum spacetime in algebraic terms
- ▶ integrating out gravitational degrees of freedom appears to yield effective NC actions with Lie-type NC structure

$$[\hat{X}^i, \hat{X}^j] = i\kappa\hbar\epsilon^{ijk}\hat{X}^k, \quad \kappa = 4\pi G \quad (\kappa\text{-Minkowski space})$$

(feasible so far only for 2+1 dimensions)

## Non-commutativity in string theory

- ▶ effective dynamics of open strings ending on D-branes: non-abelian gauge theory
- ▶ in closed string background  $B$ -field  $\rightarrow$  non-commutative gauge theory



string endpoints become non-commutative:

$$\langle f_1(x(\tau_1)) \dots f_n(x(\tau_n)) \rangle = \int dx f_1 \star \dots \star f_n, \quad \tau_1 < \dots < \tau_n$$

with star product  $\star$  depending on  $B$ .

# Spacetime non-commutativity

## Star product

General  $x$ -dependent NC structure:

$$\begin{aligned} f \star g = & f \cdot g + \frac{i}{2} \sum \theta^{ij} \partial_i f \cdot \partial_j g - \frac{\hbar^2}{4} \sum \theta^{ij} \theta^{kl} \partial_i \partial_k f \cdot \partial_j \partial_l g \\ & - \frac{\hbar^2}{6} \left( \sum \theta^{ij} \partial_j \theta^{kl} \partial_i \partial_k f \cdot \partial_l g - \partial_k f \cdot \partial_i \partial_l g \right) + \dots \end{aligned}$$

on coordinates:

$$[x^i \star, x^j] = i\theta^{ij}$$

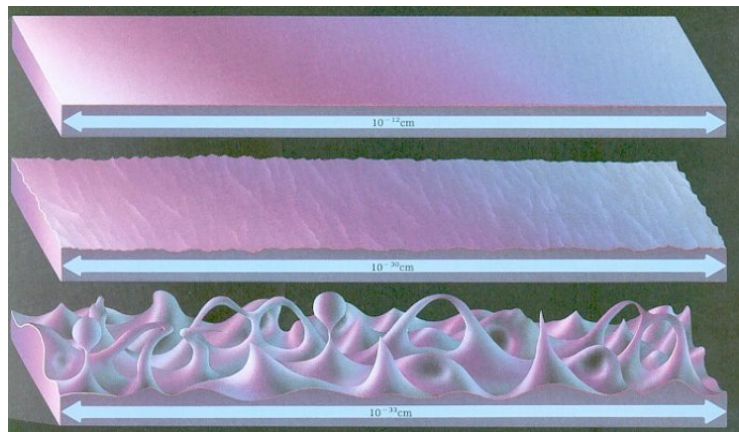
# Spacetime non-commutativity

How does a quantum space look like?



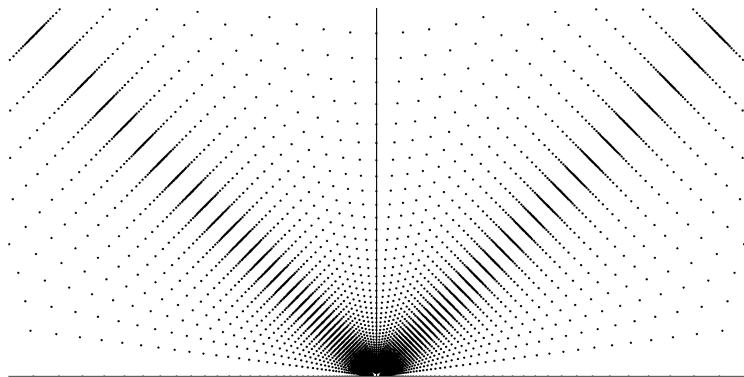
# Spacetime non-commutativity

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# Spacetime non-commutativity

How does a quantum space look like?



(q-Minkowski space; Cerchiai, Wess 1998)

## Quantum/Noncommutative Spacetime

- ▶ model of quantum geometry, spacetime uncertainty

## Particle physics on non-commutative spaces

- ▶ construction of QFT on quantum spacetime
- ▶ experimental signatures?

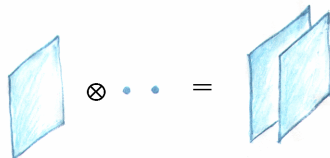
## Features

- ▶ controlled Lorentz violation, non-locality, UV/IR mixing, mixing of internal & spacetime symmetries

## Non-commutative Standard Model I

Connes, Lott; Madore (+ many others)

- ▶ spacetime augmented by a discrete space

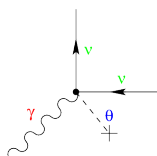
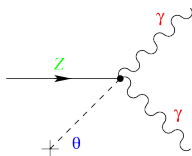


- ▶ Higgs field is NC gauge potential in the discrete direction
- ▶ beautiful geometrical interpretation of the full SM
- ▶ unlucky prediction of the Higgs mass

## Non-commutative Standard Model II

Calmet, Jurco, PS, Wess, Wohlgemant (+ many others)

- ▶ deformed spacetime with  $\star$ -product
- ▶ particle content and structure group of undeformed SM
- ▶ enveloping algebra formalism, SW maps
- ▶ many new SM forbidden interactions



## Star product and Seiberg-Witten map

Star product:

$$f \star g = fg + \frac{1}{2} \theta^{\mu\nu} \partial_\mu f \partial_\nu g + \dots$$

Corresponding expansion of fields via SW map:

$$\begin{aligned}\widehat{A}_\mu[A, \theta] &= A_\mu + \frac{1}{4} \theta^{\xi\nu} \{A_\nu, \partial_\xi A_\mu + F_{\xi\mu}\} + \dots \\ \widehat{\Psi}[\Psi, A, \theta] &= \Psi + \frac{1}{2} \theta^{\mu\nu} A_\nu \partial_\mu \Psi + \frac{1}{4} \theta^{\mu\nu} \partial_\mu A_\nu \Psi + \dots \\ \widehat{\Lambda}[\Lambda, A, \theta] &= \Lambda + \frac{1}{4} \theta^{\xi\nu} \{A_\nu, \partial_\xi \Lambda\} + \dots\end{aligned}$$

## Deformed action

Matrix multiplication is augmented by  $\star$ -products.

e.g.: noncommutative Yang-Mills-Dirac action:

$$\hat{S} = \int d^4x \left( -\frac{1}{4} \text{Tr}(\hat{F}_{\mu\nu} \star \hat{F}^{\mu\nu}) + \bar{\hat{\Psi}} \star i \hat{\not{D}} \hat{\Psi} \right)$$

with NC field strength

$$\hat{F}_{\mu\nu} = \partial_\mu \hat{A}_\nu - \partial_\nu \hat{A}_\mu - i[\hat{A}_\mu \star \hat{A}_\nu]$$

## Non-perturbative effects at the quantum level

Beta function of U(N) NC Yang-Mills theory:

$$\beta(g^2) = \frac{\partial g^2}{\partial \ln \Lambda} = -\frac{22}{3} \frac{g^4 N^2}{8\pi^2}$$

like ordinary SU(N) gauge theory – but holds also for  $N = 1$

This U(1) NC gauge theory is asymptotically free with strong coupling effects at large distance scales.



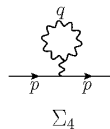
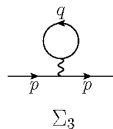
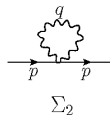
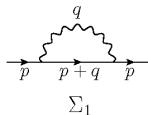
# Particle physics on non-commutative spaces

## Photon neutrino interaction

neutral particles can interact with photons via  $\star$ -commutator

$$D_\mu \Psi = \partial_\mu \Psi - i[A_\mu \star, \Psi]$$

1-loop contributions to neutrino self-energy:



## Superluminal neutrinos?

deformed dispersion relation

$$p^2 \sim \underbrace{\left( \left( \frac{8\pi^2}{e^2} - 1 \right) \pm 2 \left( \frac{8\pi^2}{e^2} - 1 \right)^{\frac{1}{2}} \right)}_{\text{constant } k, \text{ independent of } |\theta|} \cdot p_r^2,$$

direction dependent neutrino velocity

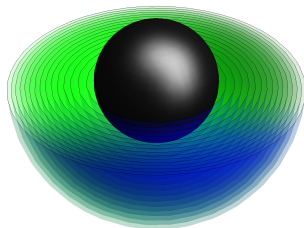
$$E^2 = |\vec{p}|^2 c^2 (1 + k \sin^2 \theta).$$

*Horvat, Ilakovac, PS, Trampetic, You: arXiv:1111.4951v1 [hep-th]*

## Non-commutative gravity

and exact solutions

- ▶ Gravity on non-commutative spacetime
- ▶ Fuzzy black hole solutions



## Motivation

### Quantum black holes

- ▶ nice theoretical laboratory for physics beyond QFT/GR
- ▶ information paradox, entropy, holography, singularities, ...

## Major obstacle

- ▶ The existence of a fundamental length scale is a priori incompatible with spacetime symmetries
- ⇒ The symmetry (Hopf algebra) must be deformed

# Non-commutative gravity

## Drinfel'd twisted symmetry:

$$\mathcal{F} = \exp \left( -\frac{i}{2} \theta^{ab} V_a \otimes V_b \right) , \quad [V_a, V_b] = 0$$

$$\Delta_\star(f) = \mathcal{F} \Delta(f) \mathcal{F}^{-1} \quad f \star g = \bar{\mathcal{F}}(f \otimes g)$$

## Twisted tensor calculus

Two simple rules:

- ▶ The transformation of individual tensors is not deformed
- ▶ Tensors must be  $\star$ -multiplied

Twisted quantization, Hopf algebra symmetry in string theory:  
Asakawa, Watamura (Tohoku University)

→ afternoon session: T. Asakawa “NC solitons of gravity”

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## Deformed Einstein equations

with “non-commutative sources”  $T_{\mu\nu}$

$$R_{\mu\nu}(G, \star) = T_{\mu\nu} - \frac{1}{2} G_{\mu\nu} T$$

## Solutions?

Solution = mutually compatible pair:

- ▶ algebra (twist)  $\star$
- ▶ metric  $G_{\mu\nu}$

# Non-commutative black hole

## Exact solution with rotational symmetry

Star product (twist) for tensors:

$$V \star W = VW + \sum_{n=1}^{\infty} C_n\left(\frac{\lambda}{\rho}\right) \mathcal{L}_{\xi_+}^n V \mathcal{L}_{\xi_-}^n W$$

left invariant Killing vector fields:  $\xi_{\pm} = \xi_1 \pm i\xi_2$

commutative radius:  $\rho^2 = x^2 + y^2 + z^2$

$$\begin{aligned} C_n\left(\frac{\lambda}{\rho}\right) &= B\left(n, \frac{\rho}{\lambda}\right) \\ &= \frac{\lambda^n}{n! \rho(\rho - \lambda)(\rho - 2\lambda) \cdots (\rho - (n-1)\lambda)} \end{aligned}$$



# Non-commutative black hole

## Metric

in isotropic coordinates

$$ds^2 = - \left( 1 - \frac{a}{\rho} \right) dt^2 + \frac{r^2}{\rho^2} (dx^2 + dy^2 + dz^2)$$

with

$$r = (\rho + a/4)^2 / \rho, \quad a = 2M$$

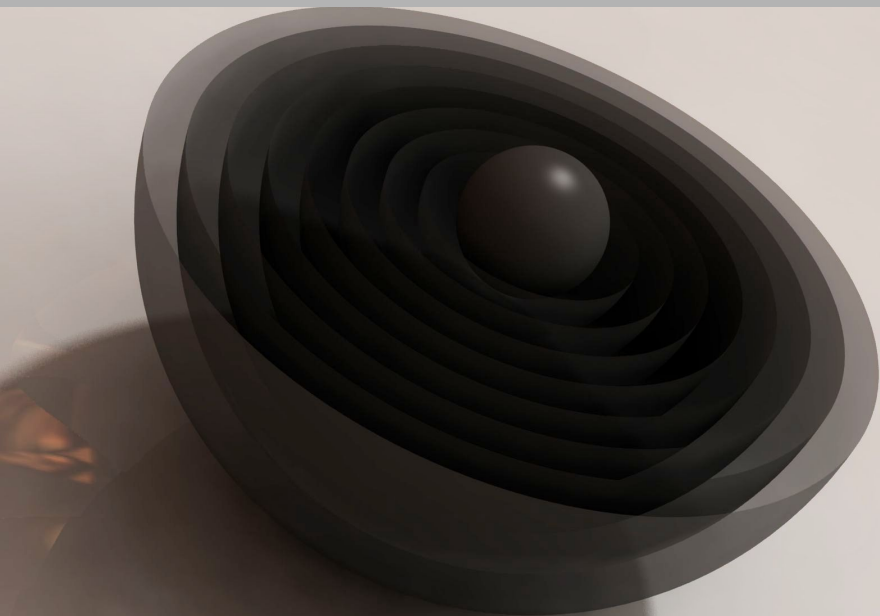
$$\rho^2 = g_{ij} x^i x^j = x^2 + y^2 + z^2$$

$$[x_i, x_j] = 2i\lambda \epsilon_{ijk} x_k$$

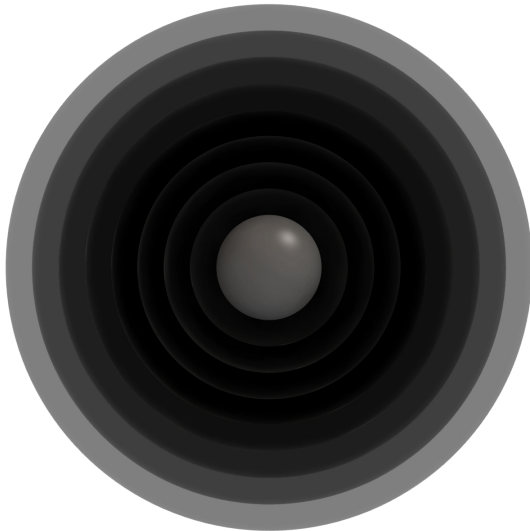
$\Rightarrow$  quantized, quasi 2-dimensional “onion”-spacetime:

$$\boxed{\rho = 2j\lambda = n\lambda; \quad n = 0, 1, 2, \dots}$$

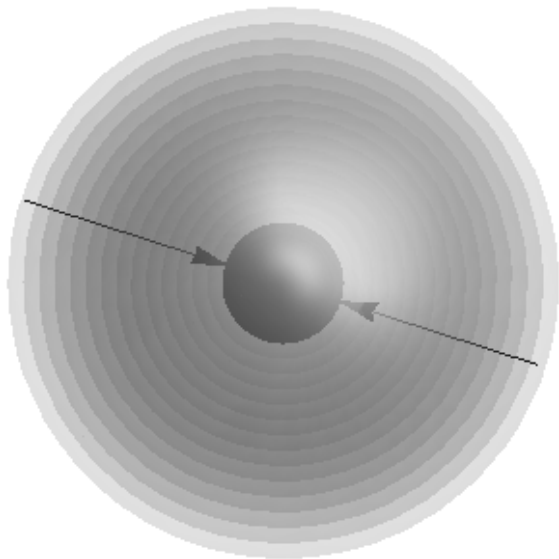
# Non-commutative black hole



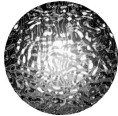
# Non-commutative black hole



# Non-commutative black hole



# Non-commutative black hole



$$L_2(\mathbb{R}^3) \rightarrow L_2(S_2)$$

We find the following surprising result:

States describing events in 3 dim NC bulk are equivalent to wave functions on a sphere (minus a fuzzy sphere)

$$\mathcal{H} = \bigoplus_{n>N} \mathbb{C}^{n+1} = \bigoplus \mathbb{C}^{n+1} - \mathcal{H}_{\text{hidden}}$$

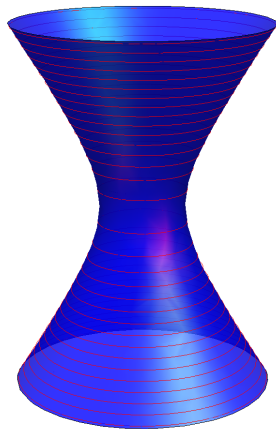
- ▶ NC Schwarzschild solution = “Fuzzy Black Hole”
- ▶ Holographic behavior appears quite naturally

$$\text{NC: bulk (3D)} \rightarrow \text{surface (2D)}$$

Heuristically:

- (1) coordinates are no longer independent:  $z \sim [x, y]$
- (2) number of commuting operators = two

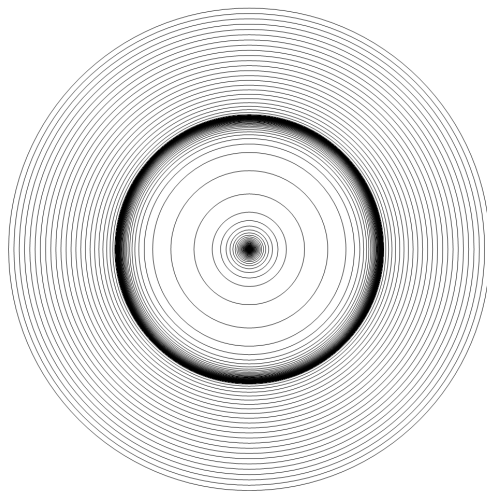
# Inside NC black hole



two sequences of fuzzy spheres with limit points  
at  $r = 0$  (singularity) and  $r = a$  (horizon)

# Non-commutative black hole

Fuzzy black hole inside and outside, in one figure:





## Generalized coherent state ( $SU(2)$ , spin $j$ representation)

$$|\Omega\rangle = \mathcal{R}_\Omega |j, j\rangle, \quad \mathcal{R}_\Omega \in SU(2)/U(1); \quad (2j+1) \int \frac{d\Omega}{4\pi} |\Omega\rangle \langle \Omega| = 1_j$$

## Star product

For  $A(\Omega) := \langle \Omega | A | \Omega \rangle$  and  $B(\Omega) := \langle \Omega | B | \Omega \rangle$  define:

$$A(\Omega) \star B(\Omega) = \langle \Omega | AB | \Omega \rangle$$

## Von Neumann entropy

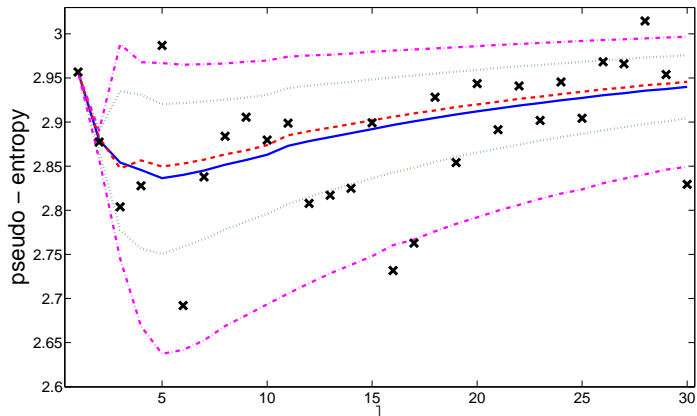
$$S_Q(\rho) = -\text{tr} \rho \ln \rho = -(2j+1) \int \frac{d\Omega}{4\pi} \rho(\Omega) \star \ln_\star \rho(\Omega)$$

Now “switch off” (or ignore) noncommutativity  $\Rightarrow$

## Wehrl entropy

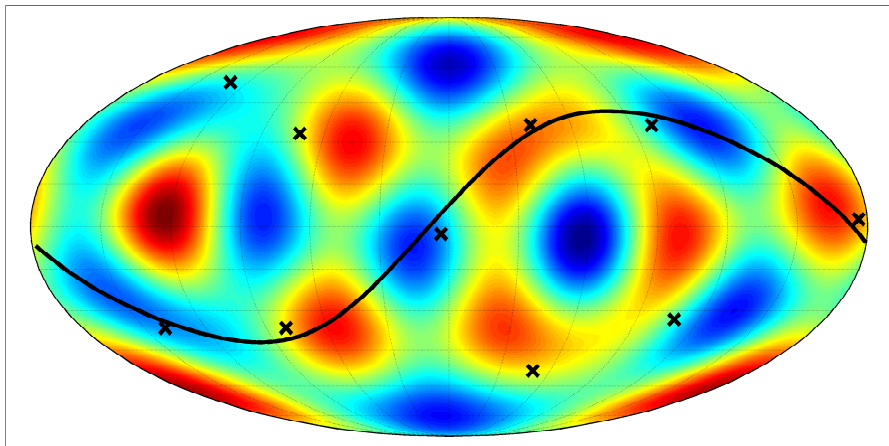
$$\begin{aligned} S_W(\rho) &= -(2j+1) \int \frac{d\Omega}{4\pi} \rho(\Omega) \ln \rho(\Omega) \\ &\geq -(2j+1) \int \frac{d\Omega}{4\pi} |\langle \Omega | \Psi \rangle|^2 \ln |\langle \Omega | \Psi \rangle|^2 \\ &> 0 \quad \text{even for pure states} \end{aligned}$$

## Coherent state analysis of CMB



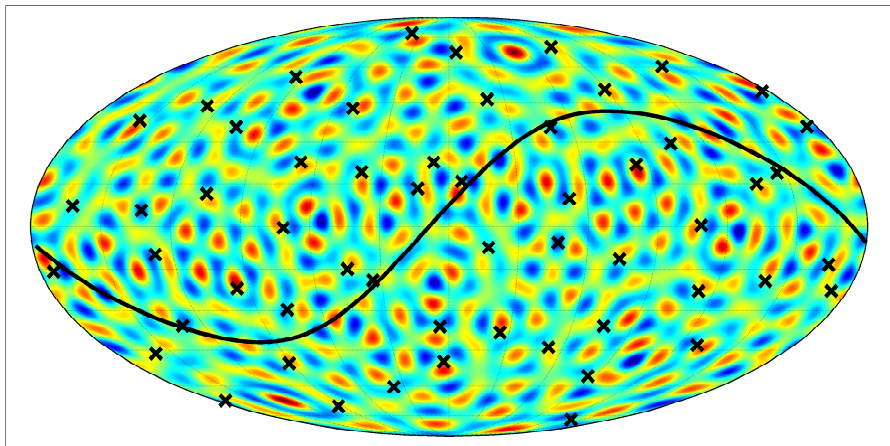
# Coherent states and CMB

CMB at  $j = 5$  with multi-pole vectors ( $\sim$  coherent states)



# Coherent states and CMB

CMB at  $j = 28$



## Beyond “non-commutative”

Recent work in M-theory reveals non-associative higher geometric structures featuring Nambu brackets:

- ▶ Basu-Harvey equations, fuzzy  $S_3$  funnels
- ▶ Bagger-Lambert action for multiple  $M2$  and  $M5$ -branes

→ afternoon session: M. Sato “3-algebra Model of M-Theory”

## Nambu mechanics

multi-Hamiltonian dynamics with generalized Poisson brackets

e.g. Euler's equations for the spinning top :

$$\frac{d}{dt}L_i = \{L_i, \frac{\vec{L}^2}{2}, T\} \quad i = 1, 2, 3$$

with angular momenta  $L_1, L_2, L_3$ , kinetic energy  $T = \sum \frac{L_i^2}{2I_i}$  and

Nambu-Poisson bracket

$$\{f, g, h\} \propto \det \left[ \frac{\partial(f, g, h)}{\partial(L_1, L_2, L_3)} \right] = \epsilon^{ijk} \partial_i f \partial_j g \partial_k h$$

## Nambu-Poisson (NP) bracket

more generally:

$$\{f, h_1, \dots, h_p\} = \Pi^{i_1 \dots i_p}(x) \partial_{i_1} f \partial_{i_2} h_1 \cdots \partial_{i_p} h_p$$

+ Fundamental Identity (FI)

$$\begin{aligned} \{\{f_0, \dots, f_p\}, h_1, \dots, h_p\} &= \{\{f_0, h_1, \dots, h_p\}, f_1, \dots, f_p\} + \dots \\ &\dots + \{f_0, \dots, f_{p-1}, \{f_p, h_1, \dots, h_p\}\} \end{aligned}$$



## Nambu-Dirac-Born-Infeld action

(B Jurco & PS 2012)

commutative  $\leftrightarrow$  non-commutative symmetry implies

$$\begin{aligned} S_{DBI} &= \int d^n x \frac{1}{g_m} \det^{\frac{p}{2(p+1)}} [g] \det^{\frac{1}{2(p+1)}} [g + (B + F) \tilde{g}^{-1} (B + F)^T] \\ &= \int d^n x \frac{1}{G_m} \det^{\frac{p}{2(p+1)}} [\hat{G}] \det^{\frac{1}{2(p+1)}} [\hat{G} + (\hat{\Phi} + \hat{F}) \hat{G}^{-1} (\hat{\Phi} + \hat{F})^T] \end{aligned}$$

This action interpolates between early proposals based on supersymmetry and more recent work featuring higher geometric structures.



Thank you for listening!